Convective overshoot in Cepheid Variables

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Cepheid variable stars are known to pulse at regular periodic intervals. These pulsations can be mathematically described in terms of a pulsation constant, Q. The relationship between this constant and the convective overshoot within a star may dictate the size of the star's helium core; and therefore its future. To test this theory, we are evolving a theoretical grid of Cepheids using MESA software, and analyzing the affects of different overshoot values using GYRE.